

Unsupervised dimensionality reduction of hyperspectral images using representations of reflectance spectra

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Abstract: Unsupervised feature selection (UFS) is a standard approach to reduce the dimensionality of hyperspectral images (HSIs). The main idea in UFS is to define a similarity metric, and select the features minimising the metric to reduce the data redundancy. In this paper, we proposed a novel criterion for unsupervised dimensionality reduction based on the representation of spectral reflectance to capture dominant reflectance variations. Since capturing all the spectral information from an entire hyperspectral dataset is a time-consuming process, we proposed a greedy search algorithm for spectral representation (GSSR). It divides the spectrum into spectral regions with less spectral variations and merges them. This algorithm, similar to FS techniques, preserves the original data from being distorted or compromised by a transformation. We compared the GSSR algorithm with well-known existing algorithms in different experiments using various datasets. Comparison with the best approximation to represent single spectra as well as entire hyperspectral scene revealed that spectral representation is almost the same. The difference between the best spectral representation and the ones provided by GSSR is less than 0.01%; while on average, GSSR is about 660 times faster to represent single spectra and 37 times faster for a complete hyperspectral scene. Five well-known unsupervised dimensionality reduction methods were also implemented and used for comparison analysis. Based on the image classification accuracy over two hyperspectral datasets, the spectral features identified by the proposed criterion improved the classification accuracy as well.

Keywords: Hyperspectral, spectral representation, unsupervised feature selection, dimensionality reduction, greedy search, image classification.

32 **1. Introduction**

33 Hyperspectral imagers, also termed imaging spectrometers, capture reflected radiance in an image
34 form, where every pixel in the image contains detailed spectral information in hundreds of adjacent
35 narrow spectral channels. Unlike multispectral sensors, with three to ten spectral bands,
36 hyperspectral sensors offer better potential for recognizing particular spectral properties
37 (Manolakis, Marden, and Shaw 2003; Shaw and Burke 2003), such as absorption bands in minerals
38 (Ben-Dor et al. 2008) or the leaf pigment content in vegetation types (Sims and Gamon 2002). On
39 the other hand, the analysis of hyperspectral data may be very challenging because well-known
40 image analysis algorithms are not easily extendable from the multispectral into the hyperspectral
41 data. Classification of hyperspectral images, for example, is often based on notions of distance in
42 the feature space, as in “minimum distance,” “minimum Mahalanobis distance,” and “k-Nearest
43 Neighbor” classifiers or in variants of k-means clustering algorithms (Gorte 1998). Part of the
44 challenge is that large sets of parameters usually are needed to describe the high-dimensional
45 statistical distributions of attributes. To have a reliable estimation of these parameters, a large
46 number of training samples is indispensable (Hughes 1968). Furthermore, the interpretation of
47 distance metrics in high dimensional spaces is not straightforward, but instead highly unintuitive
48 (Jain and Waller 1978; Jimenez and Landgrebe 1998; Durrant and Kaban 2009; Jia, Kuo, and
49 Crawford 2013).

50 A solution to the problems raised by the hyper-dimensionality is to reduce the
51 dimensionality while retaining the information required for various applications. In general,
52 dimensionality reduction (DR) is the process of reducing the number of random variables under
53 considerations. DR is categorized into two groups of feature extraction (FE) and feature selection
54 (FS) methods. FE transforms the data into a new data space based on particular criteria (Kumar,

55 Ghosh, and Crawford 2001; Jimenez-Rodriguez, Arzuaga-Cruz, and Velez-Reyes 2007). Principal
56 Component Analysis (PCA), as a classical and well-known method, eliminates the linear
57 dependency or correlation between the components (new features) in the new feature space
58 (Kaewpijit, Le Moigne, and El-Ghazawi 2002). Zhang et.al (2018) adopted manifold learning and
59 structure sparse learning algorithms to project the spectral and spatial feature into a lower-
60 dimensional subspace (Zhang et al. 2018). Recently, the low-rank matrix factorization techniques
61 showed good potential for FE as well (Zhang et al. 2019).

62 The FE techniques might have better discriminating potential between the classes in a
63 scene than the FS methods (Zaatour, Bouzidi, and Zagrouba 2017; Hira and Gillies 2015), but the
64 main issue with FE is the loss of some critical and crucial information. Since the original data are
65 no longer represented in the new data space, the information might have been compromised or
66 distorted by the transformation. FS approaches, on the other hand, have the advantage of preserving
67 the original information which is essential to analyse the spectral properties of observed materials
68 (Chein and Su 2006; Martinez-Uso et al. 2007; Carmona et al. 2011; Jia et al. 2014). These
69 techniques, also called band/channel selection, select subsets from original channels and are usually
70 preferable for analysing hyperspectral data.

71 DR can be applied using both supervised and unsupervised strategies. Having labelled
72 information, i.e., a priori knowledge about land covers in a scene paves the way for supervised DR.
73 In other words, by selecting image samples for each class, a supervised DR algorithm provides a
74 class-specific feature set. An example is the selection of channels maximizing the discrimination
75 between given classes in the feature space (Huang and He 2005; Yang et al. 2011; Hosseini Aria,
76 Menenti, and Gorte 2017). Contrary to supervised algorithms, unsupervised DR or unsupervised
77 FS (UFS) techniques do not require any a priori information. Consequently, all the pixels in an
78 image are considered for analysis. These methods are usually preferable for hyperspectral images

79 lacking the availability of labelled information (Du and Yang 2008; Cariou, Chehdi, and Le Moan
80 2011; Jia et al. 2012), which is the main objective of this study.

81 Different criteria can be applied to obtain features from a given dataset. A frequently used
82 criterion in UFS is to define a similarity metric between the spectral channels and then select those
83 channels with minimum similarity (Martinez-Uso et al. 2007; Du and Yang 2008; Cariou, Chehdi,
84 and Le Moan 2011; Jia et al. 2012). The similarity is considered as the amount of the dependent
85 information between features as well (Mitra, Murthy, and Pal 2002). The more similar the features,
86 the more the dependent. Usually, in hyperspectral data, narrow adjacent spectral channels are
87 highly correlated. As a result, the data suffer from redundancies. Therefore, by selecting less
88 dependent spectral channels, the redundant information will be minimized, and consequently, the
89 dimensionality reduces.

90 In this paper, unlike classical approaches using similarity metrics, the identification of the
91 spectral features, which accurately represent the spectral reflectance, was applied as an
92 unsupervised dimensionality reduction criterion. It means we developed an algorithm to obtain the
93 most dominant variations of spectral signals of a hyperspectral scene, which can be indicators to
94 distinguish different land covers and targets in a scene. For this purpose, instead of selecting
95 individual channels, the adjacent spectral channels were categorized based on their spectral
96 variations and then averaged; since having wider spectral bands provide more accurate image
97 classifications (Hosseini Aria, Menenti, and Gorte 2017). Hereafter, the spectral features from the
98 original hyperspectral data are called “channels,” and the ones made by averaging the neighbouring
99 channels are named “bands.”

100 To achieve the objective, the spectral signal sampled by an imaging spectrometer is
101 represented by a few spectral bands approximating the spectra with a required representation
102 accuracy; i.e., the difference between an original spectrum and the approximated one is low enough

103 to recognize a specific target using its representation spectrum (Price 1994; Jensen and Solberg
104 2007). By doing this, the most relevant spectral properties of all pixels in an image, e.g., absorption
105 features would be preserved for further analyses. When the spectral properties of pixels in an image
106 are accurately identified, they could be classified correctly with a low number of features.
107 Therefore, in this approach, a spectral band configuration is identified while minimizing the loss
108 in accuracy of representation. One of the challenges here is to find a spectral configuration, i.e. the
109 spectral locations of boundaries between spectral bands which can accurately represent all the
110 pixels in a scene. This process is mainly a very time-consuming process in unsupervised scenarios
111 since all the image pixels have to be considered for analysis. Employing a greedy search algorithm
112 (Bendall and Margot 2006; Cormen 2009) makes a locally optimal choice at each iteration and
113 provides spectral representations in a faster and more efficient fashion. Therefore, we propose an
114 algorithm called GSSR (Greedy Search for Spectral Representation) to represent the spectra and
115 evaluated it by comparisons with well-known existing algorithms for the same proposes in different
116 experiments.

117 The paper is organized as follows. Section 2 reviews the criteria frequently utilized in UFS,
118 the methods applying them to HSIs, and the algorithms applied for an accurate representation of
119 spectral reflectance. Section 3 articulates the details of the proposed criterion and how it can be
120 applied to a hyperspectral scene. The characteristics of the hyperspectral datasets used to assess the
121 proposed method are given in Section 4. Section 5 describes the evaluation procedures followed
122 by the results of different experiments, including the accuracy of spectral representation and image
123 classification. Section 6 is the conclusion.

124 **2. Related Works**

125 In this section, we first present the criteria and the methods frequently used in UFS of HSIs.

126 Secondly, we review the algorithms for accurate spectral representation, since as mentioned, our
 127 proposed criterion for UFS is to identify the most dominant spectral features from the reflectance
 128 spectra of a hyperspectral scene by accurately representing the spectra.

129 *2.1. UFS Criteria and methods*

130 This section reviews the criteria mostly used in UFS and the algorithms applying them to HSIs.
 131 These criteria are usually based on similarity (or dependency) between hyperspectral features.
 132 Accordingly, the least similar spectral features have to be selected as the ones carrying less
 133 redundant information. A group of UFS criteria are obtained by calculating the similarity between
 134 just two spectral features and creating a matrix for all the features in a dataset. This matrix is
 135 symmetric. So, if \mathbf{R} is a hyperspectral image with n spectral channels, $\mathbf{R} = \{\mathbf{R}_1, \mathbf{R}_2, \dots, \mathbf{R}_n\}$, and
 136 every channel (\mathbf{R}_i) is a vector with m pixels; $\mathbf{R}_i = \{r_{i1}, r_{i2}, \dots, r_{im}\}$, where r values are defined in
 137 space Ω , i.e. $p \in \Omega$; the similarity matrix is presented as follows:

$$\mathbf{\Sigma} = \begin{bmatrix} f_1(\mathbf{R}_1, \mathbf{R}_1) & f_1(\mathbf{R}_1, \mathbf{R}_2) & \dots & f_1(\mathbf{R}_1, \mathbf{R}_n) \\ f_1(\mathbf{R}_2, \mathbf{R}_1) & f_1(\mathbf{R}_2, \mathbf{R}_2) & \dots & f_1(\mathbf{R}_2, \mathbf{R}_n) \\ \vdots & \vdots & \ddots & \vdots \\ f_1(\mathbf{R}_n, \mathbf{R}_1) & f_1(\mathbf{R}_n, \mathbf{R}_2) & \dots & f_1(\mathbf{R}_n, \mathbf{R}_n) \end{bmatrix}, \quad (1)$$

138
 139 where $\mathbf{\Sigma}$ is the similarity or dependence matrix of \mathbf{R} , and $f_1(\mathbf{R}_i, \mathbf{R}_j)$ is the value of the dependence
 140 of the named variables. In this group, a specified search strategy is applied to the matrix and selects
 141 the channels with minimum similarity to the other channels in the dataset (Gu and Zhang 2003;
 142 Martinez-Uso et al. 2007; Qian, Yao, and Jia 2009; Jihao, Yisong, and Zhanjie 2010; Cariou,
 143 Chehdi, and Le Moan 2011; Jia et al. 2012).

144
 145 The other group of UFS criteria can be calculated as a unique score without making a
 146 similarity matrix; i.e. the similarity or dependence score is not obtained just based on two features,

147 but more than two features. So, there is no need to make a similarity matrix. Given a set of features,
 148 the score can indicate the amount of dependent or independent information of a feature in a set. In
 149 both cases, a model taking into account multiple variables applies to a hyperspectral dataset to
 150 calculate the score. So, in a general way, the score of a channel in a hyperspectral dataset can be
 151 obtained as follows:

$$D_{R_i} = f_2(\mathbf{R}_1, \mathbf{R}_2, \dots, \mathbf{R}_n), \quad (2)$$

152 where D_{R_i} is the score of the channel \mathbf{R}_i which is calculated by the model f_2 taking into account
 153 multiple channels. A list of UFS criteria mostly used and the methods of applying them are
 154 presented in the next subsections.

155 2.1.1. *Correlation Coefficient*

156 The Pearson coefficient of correlation was utilized in (Gu and Zhang 2003; Jihao, Yisong, and
 157 Zhanjie 2010) to automatically subspace hyperspectral data in an unsupervised manner. After
 158 making the correlation coefficient matrix, the authors used the local minimum of the correlation
 159 coefficient between adjacent channels of the datasets to partition the spectral channels.

160 2.1.2. *Mutual Information*

161 Mutual information (I) is a quantitative measurement of the amount of shared information between
 162 two random variables. Despite the correlation coefficient, it takes into account both linear and non-
 163 linear dependencies (Dionisio, Menezes, and Mendes 2004). Less mutual information between two
 164 random variables indicates more of uncertainty. As a result, zero, as the minimum value of the
 165 metric, means the variables are not dependent at all. It is a dimensionless quantity, generally, with
 166 units of bits (logarithms of base 2) (Cover and Thomas 2006):

$$I(\mathbf{R}_i, \mathbf{R}_j) = \sum_{\mathbf{r}_i \in \Omega} \sum_{\mathbf{r}_j \in \Omega} p(\mathbf{r}_i, \mathbf{r}_j) \log \frac{p(\mathbf{r}_i, \mathbf{r}_j)}{p(\mathbf{r}_i)p(\mathbf{r}_j)}, \quad (3)$$

167 where $p(\mathbf{r}_i, \mathbf{r}_j)$ is the joint probability distribution function of \mathbf{R}_i and \mathbf{R}_j , and $p(\mathbf{r}_i)$ and $p(\mathbf{r}_j)$ are
 168 the marginal probability distribution function of them.

169 The mutual information measure was utilized for UFS to cluster spectral channels with
 170 minimum-shared information by a recursive binary search algorithm (Cariou, Chehdi, and Le Moan
 171 2011). Martínez-Usó et al. (2007) normalized the mutual information metric and converted it into
 172 a dissimilarity metric between two channels (Martinez-Usó et al. 2007). By building a symmetric
 173 dissimilarity matrix for the entire hyperspectral dataset, a hierarchical clustering process (Jain and
 174 Dubes 1988) was applied, to form clusters of channels as similar as possible within each cluster.
 175 After obtaining k -desired clusters, a channel was selected by using a weighting method to provide
 176 the best representative channel predicting the information content of the other channels in each
 177 cluster. Eventually, it selects channels with minimum shared information. This algorithm used
 178 Ward's linkage method (Ward 1963) in hierarchical clustering, so it is named WaLuMI (Ward's
 179 Linkage strategy using Mutual Information). The advantage of this method is that it is not a ranking
 180 or incremental method that selects channels taking into account the previously selected channels,
 181 i.e., k selected channels in the clustering-based strategy are not equal to the $k-1$ selected channels
 182 plus another relevant channel.

183 2.1.3. *Kullback-Liebler Divergence*

184 This metric is also based on the information theory, which was applied for UFS of hyperspectral
 185 images (Martinez-Usó et al. 2007; Qian, Yao, and Jia 2009). The metric was considered as a
 186 dissimilarity distance between two probability distributions and interpreted as the cost of using one
 187 of the distributions instead of the other one (Martinez-Usó et al. 2007). Martínez-Usó et al. (2007)

188 applied an algorithm similar to WaLuMI by replacing the normalized mutual information metric
189 with the Kullback-Liebler criterion, and they named it WaLuDi (Ward's Linkage strategy using
190 Divergence).

191 2.1.4. *Euclidean Distance*

192 The negative Euclidean distance was used as a similarity measure in an affinity propagation-based
193 channel selection algorithm. It is used to indicate how well a spectral channel represents other
194 channels by making a similarity matrix. The channels are, then, clustered based on their similarities
195 (Jia et al. 2012).

196 2.1.5. *Dependent Information Metric*

197 Sotoca. et al. (2007) defined a metric to obtain the dependent information of a set of random
198 spectral channels. The set can have more than two variables, and therefore, there is no need to make
199 a dependence matrix of pairs of variables (Sotoca, Pla, and Sanchez 2007). The metric measures
200 the dependent information of a set of channels by employing the joint entropy and the conditional
201 entropy. Applying the metric in a greedy search algorithm, the authors selected channels having
202 the minimum-shared information (Sotoca, Pla, and Sanchez 2007).

203 2.1.6. *Linear Prediction*

204 Du et al. (2007) applied two similarity-based endmember extraction algorithms to select spectral
205 channels in an unsupervised manner (Plaza et al. 2004; Du and Yang 2007, 2008). They searched
206 the most distinctive channels based on linear unmixing methods. These methods model an HSI
207 pixel as a linear mixture of a set of finite image endmembers and select the most distinctive
208 channels in a sequential forward selection searching strategy (Du, Ren, and Chang 2003). They can

209 jointly evaluate the similarity between a given channel and multiple channels.

210 The first one, linear prediction (LP) criterion, makes a linear estimation of the selected
 211 channels and searches for the most dissimilar one. To find a channel that is the most dissimilar to
 212 channel \mathbf{R}_1 and \mathbf{R}_2 , one can first estimate \mathbf{R}' as the linear prediction of the two channels by solving
 213 the following linear model using the least-squares solution:

$$\alpha_0 + \alpha_1 \mathbf{R}_1 + \alpha_2 \mathbf{R}_2 = \mathbf{R}', \quad (4)$$

214 where α_0 , α_1 , and α_2 are the parameters that can minimize the LP error. Then they searched for a
 215 channel having the maximum dissimilarity with the estimated channel, i.e., having the maximum
 216 $e = \|\mathbf{R}' - \mathbf{R}_i\|$, where $\|\cdot\|$ can be the Euclidean norm of the vector.

217 2.1.7. *Orthogonal Subspace Projection*

218 Du et al. (2007) used orthogonal subspace projection (OSP) as the second criterion to map a channel
 219 onto an orthogonal subspace (Du and Yang 2007). With this criterion, first, an orthogonal subspace
 220 based on the pre-selected channels is constructed. Then, each channel is projected into the
 221 orthogonal space. The channel yielding the maximum orthogonal component is considered as the
 222 most dissimilar one to the pre-selected channels.

223 2.1.8. *Independent Component Analysis*

224 Independent component analysis (ICA) is a method that extracts independent source signals by
 225 searching for a linear transformation that minimizes the statistical dependence between the
 226 components (Comon 1994). Unlike PCA, ICA imposes higher-order dependence than the second-
 227 order one, so the components are not always orthogonal (Comon 1994; Hyvärinen and Oja 2000).
 228 ICA is used as a feature extraction technique in several approaches (Chiang, Chang, and Ginsberg

229 2000; Lennon et al. 2001; Robila and Varshney 2004; Liu et al. 2017), while Du et. al. (2003)
 230 presented an ICA-based method for feature selection as well to reduce the dimensionality of HSIs
 231 (Du et al. 2003). The authors weighed the spectral channels using the independent components and
 232 selected those having maximum information.

233 The ICA-based model represents a hyperspectral image, \mathbf{R} , as the multiplication of an
 234 unmixing matrix, \mathbf{A} , and a number of independent components called “sources”, \mathbf{S} :

$$\mathbf{R}_{n \times m} = \mathbf{A}_{n \times c} \times \mathbf{S}_{c \times m}, \quad (5)$$

235
 236 where c is the number of source signals. Following this model, the ICA aims at deriving the best
 237 possible estimation of \mathbf{S} by approximating the unmixing matrix \mathbf{A} under some constraints. Based
 238 on the approximation of the unmixing matrix, Du et. al. (2003) calculated a mean absolute weight
 239 per spectral channel as the indicator of the information content of each channel. These weights
 240 were then sorted, and the channels with the highest weights were selected as the most informative
 241 ones.

242 ***2.2. Spectral representation***

243 There are several methods for spectral approximation and representation (Price 1975; Price 1990;
 244 Li et al. 1999; Wang et al. 2007; Huynh and Robles-Kelly 2008; Angelopoulou 2000;
 245 Angelopoulou, Molana, and Daniilidis 2001), mostly used in colourimetric sciences. Their scope
 246 is to represent spectra accurately with a limited number of samples e.g. Discrete Fourier Transform
 247 (DFT) (Agrawal, Faloutsos, and Swami 1993), Singular Value Decomposition (SVD) (Keogh et
 248 al. 2001) or Discrete Wavelet Transform (DWT) (Kahveci and Singh 2001). A common way for
 249 the approximation is the replacement of local variations in a spectrum with a constant value over a
 250 small range in wavelength. Chakrabarti et. al. (2002) presented such a technique, named Adaptive

251 Piecewise Constant Approximation (APCA), and proved that this technique yields a better
252 representation than other existing methods such as DFT and DWT for approximating signals in
253 time series analyses (Chakrabarti et al. 2002). The APCA algorithm degrades a curve into a
254 constant segment-based approximation, where the user specifies the number of segments. It
255 includes two main steps. At first, it converts the signal approximation issue into a wavelet
256 compression problem, for which there are well-known optimal solutions; and next, it converts the
257 solution back to the APCA representation and makes minor modifications. The term ‘segment’ is
258 equivalent to ‘band’ in our approach. More details on APCA can be found in (Chakrabarti et al.
259 2002). Approximating spectra by piecewise constant functions has also been used in other fields,
260 taking into account the physical characteristics of the spectra to determine the location of the
261 spectral segments (Thomson, Lue, and Bannerman 2014; Zehentbauer and Kiefer 2012).

262 Konno and Kuno (1998) proposed a method that provides the best piecewise approximation
263 (Konno and Kuno 1988). They used the maximum norm and the Euclidean norm to find the
264 approximation of a function of a single variable with less than a predefined number of constant-
265 value segments. Another study applied the Bayesian approach for piecewise smoothing of one-
266 dimensional signals (Winkler and Liebscher 2002). Later, an extension of this method was used
267 for multiple spectral curves to reduce the dimensionality of hyperspectral scenes (Jensen and
268 Solberg 2007). The goal of the last approach was also to partition the spectra of a hyperspectral
269 scene into a fixed number of contiguous intervals with fixed intensities using the piecewise constant
270 function approximations (PCFA) algorithm. The intensity in a spectral band is the mean value of
271 the signal in its constitutive channels per pixel. Considering the number of bands, the algorithm
272 examines all the possible spectral locations for the breakpoints and finds the best approximation
273 having the lowest error of representation.

274 3.Spectral Representation Criterion

275 As reviewed, the criteria used in UFS algorithms applied to HSIs are based on defining a similarity
276 metric between spectral features and selecting the ones having minimum similarity with other
277 features. Following this procedure, the redundant information is decreasing while the
278 dimensionality of the original dataset is reducing. In this approach, we suggest a criterion for UFS
279 with regards to capturing the most relevant spectral information concerning all the pixels in a
280 hyperspectral image. It means all the spectral reflectance of an entire scene are approximated in a
281 way that the approximated spectra represent the original ones accurately.

282 For this purpose, we partition the reflectance signals of a hyperspectral scene into a
283 predefined number of adjacent bands with fixed intensities. Figure 1 schematically illustrates the
284 idea for a spectral signature with 195 channels, which is divided into six bands. In this figure, the
285 blue line is the representation (approximation) of the original spectral signal (the red line). In fact,
286 using this method, the channels are transformed into a new feature space but the transformation is
287 in a way that the relations between the reflectance spectra and their wavelength are retained, similar
288 to the FS techniques.

289 *Figure 1 ---- >>>*

290 We use the square error between the original reflectance spectrum (r) and the approximated
291 (a) one to identify the error of representation. Since the intensity is a constant value in every
292 spectral band of the representation; to minimize the error, the intensity of a spectral band should
293 be the mean value of its constitutive channels. So, the root means square error (RMSE) is used to
294 calculate the error of spectral representation. For a complete hyperspectral scene, the following
295 procedure is performed.

297 let the hyperspectral dataset defined in Section 2 be divided into $k+1$ bands where $k \leq n$
 298 and k is the number of breakpoints. The set of the spectral locations of breakpoint is $\mathbf{S} =$
 299 $\{\mathbf{s}_0, \mathbf{s}_1, \dots, \mathbf{s}_k, \mathbf{s}_{k+1}\}$; where $\mathbf{s}_0 = 0$, $\mathbf{s}_{k+1} = n$, and $\mathbf{s}_1, \dots, \mathbf{s}_k$ indicate the channel numbers in an
 300 ascending order where the breakpoints are placed after them. Therefore, a new band set with m
 301 pixels in each band is $\mathbf{A} = \{\mathbf{A}_1, \mathbf{A}_2, \dots, \mathbf{A}_k\}$ where

$$\mathbf{A}_i = \frac{\sum_{s_{i-1} < t \leq s_i} \mathbf{R}_t}{s_i - s_{i-1}}. \quad (6)$$

303
 304 To compute the error of representations for the entire scene, first, the reduced spectral
 305 configuration is expanded at each pixel back into the original channel configuration. Then, the
 306 values of each band are duplicated in the adjacent channels covered by the band. Next, the
 307 reconstructed and full spectra per pixel are compared by computing the RMSE between the two
 308 spectra. The difference between the expanded and the original spectra, then, can be calculated and
 309 averaged over all pixels to obtain the error (E_{rep}) of the representation for the entire scene:

$$E_{\text{rep}} = \frac{\sum_{j=1}^m \sqrt{\frac{1}{n} \sum_{i=1}^n (\mathbf{r}_{ij} - \mathbf{a}_{ij})^2}}{m}, \quad (7)$$

310 where \mathbf{r}_{ij} and \mathbf{a}_{ij} are the i^{th} signal value in the j^{th} pixel of the original and the approximated spectra
 311 respectively.

312 Having the representation error of different band configurations, we can select the optimal
 313 band set representing the signals with adequate accuracy. There is a huge number of combinations
 314 to select the location of the breakpoints and every band configuration gives different representation
 315 errors. Ideally, the best locations are the places where the total error (E_{rep}) is minimum. It can be
 316 achieved by an exhaustive search (Nievergelt 2000), i.e. all band configurations with the given
 317 number of bands are considered and evaluated. In a practical situation, however, the computational

318 cost for large datasets is prohibitive. This method can be used for a limited number of spectral
319 signatures (Jensen and Solberg 2007).

320 In our approach, we apply a greedy search strategy to determine the spectral location in a
321 sequential manner. This strategy was used in the FS algorithms to apply different criteria for DR
322 of hyperspectral images as well (Pudil, Novovicova, and Kittler 1994; Sotoca, Pla, and Sanchez
323 2007; Le Moan et al. 2011; Yang et al. 2011; Han, Lee, and Bien 2013; Hosseini Aria, Menenti,
324 and Gorte 2017). Using the spectral representation as a criterion, we named the algorithm as Greedy
325 Search for Spectral Representation (GSSR). In this procedure, the algorithm iteratively selects a
326 spectral location of a breakpoint that appears to be the best with regards to the representation error
327 and the previously selected subset of breakpoints. The method significantly reduces the complexity
328 by progressively ranking the evaluated subset.

329 ***3.1. Greedy search for spectral representation (GSSR).***

330 Given a hyperspectral dataset, \mathbf{R} , and the number of bands, $k+1$, the algorithm scans all possible
331 spectral locations by taking into account the previously selected breakpoints to determine a new
332 breakpoint in each iteration. For every tentative breakpoint, it creates the band set (A) (Eq. (6))
333 based on the preselected breakpoints and the new one, and then calculates the representation error
334 (E_{rep}). Therefore, all possible locations for a new breakpoint are examined, and the best
335 approximation having the lowest error of representation is identified. Then, the determined
336 breakpoint is added to set S . This procedure terminates when the number of bands reaches the
337 predefined value ($k+1$). Figure 2 illustrates the flowchart of the procedure.

338 *Figure 2 ---- >>>*

339

340 The GSSR algorithm complexity for a single spectrum is of the order $O(kn)$, where k is the
341 number of breakpoints, and n is the number of channels in the original hyperspectral dataset. The
342 complexity of the metric for spectra representation is of the order $O(mn)$, where m is the number
343 of pixels. Hence, the overall computation time of the GSSR algorithm is $O(kmn^2)$.

344 In the next section, the data sets used for the assessment of the proposed algorithm are
345 presented.

346 **4. Hyperspectral Datasets**

347 The GSSR algorithm was evaluated by applying it to different hyperspectral datasets, including a
348 spectral library. We have done the necessary pre-processing steps before using the datasets,
349 including atmospheric correction and removal of the noisy channels for the scenes. The noisy
350 channels are those that do not have any signal, located at water absorption spectral regions, and the
351 ones having a low signal to noise ratio (SNR). The channels with low SNR were identified by
352 estimating the SNR using the geostatistical method described in (Curran and Dungan 1989), and
353 visual inspection. We used the following datasets for the experiments:

- 354 • A spectral library: it comprised 1365 spectra from different materials and was developed
355 by researchers at the Spectroscopy Lab, USGS, in 2007. The library is divided into six
356 chapters: 1. Minerals, 2. Mixtures, 3. Coatings, 4. Volatiles, 5. Man-Made; and 6. Plants,
357 Vegetation Communities, Mixtures with Vegetation, and Microorganisms. There is more
358 than one spectrum for many of materials since different factors have been considered for
359 the collection of the spectra such as the type of the spectrometer, the spectral resolution, the
360 purity of the materials, the grain size, the presence of other elements in the sample, etc. The
361 chapters contain 881, 138, 12, 24, 110, 200 spectra respectively. The library is used as a

362 reference for material identification in remote sensing images. The database is over 6000
363 webpages. More details of the spectral library can be found at
364 <https://speclab.cr.usgs.gov/spectral-lib.html>. We used the convolved version of the library
365 corresponding to the AVIRIS channels. After analyzing the spectral library, we found out
366 that three pairs of the spectra are the same, and they cannot be distinguished from each
367 other. These spectral pairs are from Mixtures and Plants chapters. The duplicated spectra
368 were removed. Consequently, the final number of spectra in Chapter 2 and 6 became 136
369 and 199 respectively.

370 • Moffett Field: AVIRIS has acquired this dataset in California with 224 bands. The band
371 set covers the spectrum from 365nm to 2497nm continuously with approximately 10nm-
372 wide channels. The channels located at 366-385, 1353-1433, 1811-1948, 2337-2497 nm
373 wavelength were removed due to noise and water absorption. As a result, the final dataset
374 has 177 channels (Figure 3).

375 *Figure 3 ---- >>>*

376 • Indian Pines: the scene consists of 145*145 pixels with a spatial resolution of about 20m.
377 Two-thirds of the Indian Pines scene is covered by agriculture, and one-third by forest and
378 other natural perennial vegetation (Figure 4). The ground truth available documents sixteen
379 classes, not mutually exclusive. Since three classes in the scene contain less than 50
380 samples, we do not use them for the experiments. After the atmospheric correction and the
381 removal of noisy channels, the number of channels was reduced to 178. We removed water
382 absorption channels (104-108, 150-163, and 220), noisy bands (1-4, 103, 109-111, 148-
383 149, 164-166, and 217-219), and seven channels that are spectrally overlapping channels
384 (32, 33, 95, 96, 158, 191, and 192). The Indian Pines dataset is available free of charge via

385 Purdue University website:
386 <https://engineering.purdue.edu/~biehl/MultiSpec/hyperspectral.html>.

387 *Figure 4 ---- >>>*

388 • Salinas: This scene (Figure 5) is characterized by high spatial resolution (3.7m). The area
389 covered comprises 512 lines by 217 pixels. The dataset is available at
390 http://www.ehu.es/ccwintco/index.php/Hyperspectral_Remote_Sensing_Scenes only as
391 at-sensor radiance. So, it has been atmospherically corrected, and the noisy and duplicated
392 channels have been removed. The final dataset has 190 channels. The ground-truth is also
393 available and documents 16 classes, including vegetables, bare soils, and vineyard fields,
394 which we used in the experiments.

395 *Figure 5 ---- >>>*

396 **5. Evaluation of the Proposed Method**

397 Two types of experiments were performed to evaluate the GSSR algorithm: a) single signal
398 representation and b) unsupervised dimensionality reduction of hyperspectral scenes. In both types
399 of experiments, the accuracy of representation and the running time were evaluated, while for the
400 second type of experiment, the image classification accuracy was also considered. We performed
401 the assessments by comparing the algorithm with well-known existing algorithms for the same
402 purposes.

403 ***5.1. Evaluation of the algorithm using single spectra***

404 Two experiments were performed in order to assess the GSSR algorithm for representing a single
405 spectrum. In the first experiment, the GSSR algorithm is compared with APCA and PCFA (Section

406 2.2) using various spectra. Both algorithms represent spectra with a set of constant signal value
407 segments as GSSR (Figure 1). In this experiment, the three algorithms were compared in terms of
408 the spectral representation and the running time.

409 In the second experiment, the reduced spectral configurations obtained by the algorithms
410 were evaluated for material detection, i.e. different materials were detected by comparing the
411 approximated spectra with the full spectra available in the spectral library.

412 *5.1.1. Experiment 1: single spectra representation*

413 This evaluation was performed as a benchmark to identify the error of representation of single
414 spectra using the three methods: GSSR, PCFA, APCA. At first, three dominant reflectance spectra;
415 soil, water, and vegetation, were compared, and the three algorithms were applied to represent the
416 spectra with 5, 10, and 15 bands. These reflectance spectra were obtained from the pixels with the
417 same land cover in the Moffett Field scene. Figure 6 illustrates the results, and Table 1 gives the
418 error of the estimate by different algorithms for all the spectral configurations.

419 *Figure 6 ---- >>>*

420

421 *Table 1 ---- >>>*

422

423 As expected, the PCFA algorithm gave the smallest error of estimate in all cases, while
424 APCA gave the largest error (Table 1). Interestingly, GSSR represented the spectra almost twice
425 as accurately as APCA and with an accuracy comparable with the PCFA algorithm. For example,
426 using ten bands to approximate the soil spectrum, the error of the estimate was 0.0230 with APCA,
427 0.0115 with GSSR, and 0.0096 with the PCFA algorithm. The difference in the error between
428 GSSR and PCFA is less than 0.002, and it became lower when 15 bands were used, with the

429 difference in RMSE being 0.0005 only. It is also observed (Figure 6) that the locations of the
430 breakpoints determined by the PCFA and GSSR algorithms are almost identical.

431 The GSSR and PCFA algorithm always divide the spectrum into the exact predefined
432 number of bands, while APCA does not, as shown in this experiment. This situation occurred, for
433 example, when seeking to approximate the water spectrum, with ten bands by the APCA algorithm.
434 In this case, the reduced spectral configuration had one spectral band less than the prescribed
435 number of bands, while the error of estimate would have been lower with one additional band.
436 APCA is based on the Haar wavelet transform, so the number of samples in the original signal fed
437 into the algorithm has to be a power of two. In the case that the signal does not have enough
438 samples, it is padded with zeros, and later truncated. This process sometimes may yield fewer bands
439 than expected.

440 We repeated the same experiment using more than 1000 pixels with different reflectance
441 spectra derived from the Moffett Field AVIRIS image. The pixels were chosen to sample various
442 land cover types, including different types of water, soil, vegetation, man-made features such as
443 buildings, roads, etc. The reflectance spectra were represented separately for each pixel with a
444 different number of bands starting from 5 to 30, in steps of 5. Figure.7 shows the results.

445 *Figure 7 ---- >>>*

446
447 The mean RMSEs decrease with an increasing number of bands. Similar to the previous
448 results, the APCA error was the largest one, while the error for GSSR and PCFA algorithms were
449 very similar and lower than when using APCA. The mean RMSE difference between PCFA and
450 GSSR over all the spectra samples was about 0.0003 in the unit of the spectral reflectance, i.e. in
451 [0,1], with the 5-band representation to 0.0001 with the 30-band one. The difference between GSSR
452 and PCFA representation errors decreased with an increasing number of bands.

453 We also considered the run time required to carry out the numerical experiment on the
454 dataset with more than 1000 spectra (Table.2). The time estimate is based on the implementation
455 of the algorithms on a desktop computer that has the following characteristics: Operating system:
456 Windows 7, Processor: Intel Core 2 and 16 GB RAM. The algorithms were written in IDL
457 programming language, version 8.2.

458 *Table 2 ---- >>>*

459
460 The APCA algorithm is fast. The running time was less than two seconds for all the spectral
461 configurations, while it was increasing with the number of bands for GSSR: started at less than 14
462 seconds for the 5-band representations and reached more than two and half minutes for the 30-band
463 ones. PCFA consumed much more time than the two other algorithms. In the worst case, i.e. the
464 30-band representation, the run time was more than two days to find the representations for the
465 1089 spectra, which was about 1200 times slower than GSSR. On average, GSSR ran 660 times
466 faster than PCFA. The main issue affecting the running time of the PCFA algorithm is that the
467 algorithm recursively calls itself with respect to the number of bands, and checks all the possible
468 situations. When the number of bands increases, the run time increases dramatically.

469 *5.1.2. Experiment 2: material detection using approximated spectral signatures*

470 The previous experiment showed that the GSSR algorithm yields comparable spectral
471 configurations to the best representation provided by PCFA with much shorter run time. In the
472 second experiment, we evaluated the spectral configurations obtained by GSSR by applying them
473 to material detection.

474 This experiment reveals the number of bands needed to correctly identify a target spectral
475 signature using the reduced spectral configurations derived by the algorithms GSSR and PCFA.
476 The APCA algorithm was omitted since the spectral representations provided by this algorithm are

477 not as accurate as of the representations obtained by the other algorithms. For this experiment, we
478 used the spectral library that contains different and well-defined spectra.

479 In this experiment, a spectrum from the library is selected as a “target spectrum.” Then the
480 GSSR and PCFA algorithms were applied to identify the breakpoints in such a way that is the
481 reduced spectral configuration represents the full spectrum with increasing accuracy. At each
482 iteration, the reduced spectral signature of the target spectrum was compared with all the spectra
483 in the spectral library to check whether the approximated spectrum could correctly be identified,
484 i.e. the approximated target spectrum and the full detailed one has the least difference. The iterative
485 procedure was ended when the reduced spectral configuration of the target signature had been
486 correctly identified, or the number of bands was more than 30. We used a distance-based identifier
487 and a spectral angle based identifier to measure the difference between the known and unknown
488 spectra (Kruse et al. 1993; Price 1994; Cochrane 2000). Finally, we calculated the percentage of
489 spectra correctly identified vs. the number of bands (Figure 8).

490 *Figure 8 ----- >>>*

491
492 In general, the accuracy of the distance-based identifier is higher than the angle-based one.
493 For instance, the 10-, 15-, and 20-band spectral configurations achieved correct identification of
494 materials in 97%, 99%, and 100% of cases with the distance-based identifier, while with the angle-
495 based identifier, the correct identification reached to 61%, 79%, 86% respectively. The latter
496 normalises the spectra and removes the signal intensity dependence, i.e., reflectance in this
497 experiment.

498 Using either identifier, the reduced spectral configurations obtained with PCFA and GSSR
499 gave a comparable accuracy in material detection. The spectral configurations obtained with PCFA
500 gave slightly more identifications than the ones obtained with GSSR. The difference in

501 performance between PCFA and GSSR was higher when the number of bands is small, and it
502 decreased with an increasing number of bands. However, if the spectral configuration obtained
503 with GSSR has just one band more than the PCFA configuration, the detection accuracy for GSSR
504 is higher. For example, using the angle-based identifier, the 16-band configurations obtained by
505 GSSR were correctly identified in 80.7% of the cases, while the 15-band configuration obtained
506 with PCFA was accurate in 80.6 % of cases. It should be noted that the computational cost of the
507 15-band PCFA configuration is much higher than the one of the 16-band configuration obtained
508 by GSSR. In the example mentioned, GSSR was more than 2000 times faster than PCFA.

509 Both experiments (Section 5.1.1 and 5.1.2) revealed that the spectral representation of
510 single spectra using GSSR is almost identical with the best spectral representation and has
511 comparable accuracy with it in the representation and in detecting materials using the reduced band
512 configurations. Meanwhile, GRRS provides the representations in a much faster way than
513 obtaining the best representation. In the second type of experiment, we evaluated the algorithm
514 applying to an entire hyperspectral scene.

515 *5.2. Evaluation of the algorithms using the entire scene*

516 In the second type of experiment, the GSSR algorithm was evaluated by two experiments
517 using the entire hyperspectral scenes. In this case, the spectral locations of the breakpoints must be
518 the same for all pixels to reduce the dimensionality of the image. At first, it was again compared
519 with PCFA to assess the error of representation and the running time, since PCFA provides the best
520 spectral approximations for all the spectral reflectance in a hyperspectral scene with the same
521 situation as GSSR. It means that both algorithms take the average of adjacent channels to form
522 wider spectral bands. Therefore, the band configurations, i.e., the spectral locations of breakpoints
523 over the spectrum identified by the algorithms, exert the principal influence on the spectral

524 representations; having the same spectral configurations would provide similar spectral
525 representations.

526 The second experiment was a standard methodology to compare different feature sets
527 obtained by various algorithms in an image classification procedure (Shaw and Burke 2003;
528 Martinez-Uso et al. 2007; Sotoca, Pla, and Sanchez 2007; Cariou, Chehdi, and Le Moan 2011; Jia
529 et al. 2012). A better image classification generally means that the process of assigning a label to
530 a pixel using its spectral information is more accurate, which leads to better recognition of objects
531 and land covers in the image.

532 Therefore, to validate the performance of the proposed method, we presented a comparison
533 with five other unsupervised DR algorithms by evaluating the image classification accuracy. Four
534 of them are in the FS category, and the last one is in the FE category. The FS algorithms are
535 WaLuMI, LP, OSP, and ICA-based model. Similar to GRRS, they preserve the physical
536 relationship between the selected features and their wavelength. The FE algorithm is the Principal
537 Component Analysis (PCA). We compared the GSSR algorithm with PCA since GSSR, similar to
538 PCA, transforms the data into a new feature space but without using a rotation.

539 *5.2.1. Experiment 1: HSI Spectra Representation:*

540 In this experiment, we applied the PCFA and GSSR algorithm to an entire Moffett Field
541 dataset. The dataset contains various spectral reflectance with different variations in the spectrum.
542 We computed the mean RMSE of the spectral configurations provided by the PCFA and GSSR
543 algorithms with respect to the number of bands in the reduced spectral configuration. The error is
544 the average of the spectral approximation error of all pixels in the scene (Eq. (7)).

545 It should be noted that PCFA was developed to minimize the sum of squared error (SSE)
546 as a performance metric. However, minimizing SSE concerning the approximated spectrum with
547 k predefined number of breakpoints is equivalent to minimizing RMSE with the same conditions.

548 The running time of the algorithms is also computed. Figure 9 shows the results and Table
549 3 presents more details about six band sets. The error of spectral representation using the same
550 configuration for an entire scene reveals an almost complete overlap between the two graphs
551 showing the mean representation error obtained by the PCFA and GSSR algorithm. Table 3
552 indicates that the difference between the mean errors of the two methods is about 0.0001. On the
553 other hand, the PCFA was about 37 times slower than GSSR, on average. The higher the number
554 of bands, the slower the PCFA than GSSR, as clearly illustrated in Figure 9 (left).

555 *Figure 9* ---- >>>

556

557 *Table 3* ---- >>>

558

559 Jensen et. al. (2007) applied PCFA to a set of sampled spectra of a hyperspectral scene
560 derived from the classes in a scene and identified a single spectral configuration for the spectra to
561 reduce the dimensionality of the data (Jensen and Solberg 2007). The PCFA algorithm is applicable
562 when the number of spectra is low; however, when the number increases, PCFA is slow. The
563 complexity of the algorithm is $O(kmn^3)$ (Jensen and Solberg 2007), i.e. the order of the algorithm
564 has a direct relation with the cube of the number of spectral samples multiplied by the number of
565 pixels. Therefore, if the number of pixels increases, the time consumption of the algorithm
566 drastically goes up. As a consequence, the PCFA algorithm becomes a prohibitive method in an
567 unsupervised DR situation, where applied to all pixels in a hyperspectral scene.

568 On the other hand, GSSR provides a spectral configuration for the entire scene as accurate
569 as of the best spectral configuration supplied by the PCFA algorithm but in a much faster way.

570 *5.2.2. Experiment 2: Image Classification*

571 The evaluation has been done by examining the number of features selected by the proposed and
572 reference methods vs. the classification overall accuracy using different classifiers to check the
573 relevance of the features selected. We applied the five mentioned methods to compare the results
574 with the band sets obtained by the GSSR method. These comparisons were performed using two
575 datasets: the Indian Pines and Salinas scene. We used two types of classifiers: maximum likelihood
576 classifier (MLC) and support vector machine (SVM), a parametric classifier, and a non-parametric
577 classifier, respectively. Figure 10 shows the results.

578 *Figure 10 ---- >>>*

579
580 As observed, the proposed method gives better overall accuracy than its competitors from
581 the FS category. It means that the criterion used in our approach in an unsupervised manner, i.e.,
582 extraction the most relevance spectral features by spectral reflectance representation provides
583 higher accuracy of classification than the frequently used criteria based on the similarity between
584 the spectral features in UFS techniques.

585 Comparison with PCA, an FE algorithm, shows that GSSR has better accuracy in
586 classification when using MLC while using the SVM classifier, PCA provided higher accuracy.
587 The reason is that the feature extraction techniques often have a higher potential in distinguishing
588 between different classes in a scene, leading to better accuracy in image classification. However,
589 the problem of the FE algorithms is that the critical information of the reflectance spectra can be
590 distorted. One of the main objectives of this study is to keep the critical information of the
591 reflectance spectra like the FS methods. This information, e.g. the absorption spectral features of a
592 specific target, is of interest to a wide range of HSI users. The GSSR algorithm, while retaining the

593 key spectral information, classified more accurately than PCA when using MLC. In addition, it
594 obtained a better result than its competitors in the UFS category as well.

595 Using SVM, WaLuMI sometimes provided channel sets with comparable classification
596 accuracy to GSSR. The channel selection algorithms based on ICA and the linear unmixing
597 methods (LP and OSP) mostly gave less accurate results than GSSR.

598 **6. Conclusions**

599 This study showed the advantage of applying the representation of reflectance spectra of HSIs as
600 the criterion to the unsupervised dimensionality reduction purpose. The typical rules applied in
601 unsupervised feature selection techniques are based on finding the most dissimilar spectral
602 channels, while the proposed criterion focuses on the extraction of the most spectral variations
603 from the spectral reflectance. Since obtaining accurate spectral representations for all the pixels in
604 a scene; i.e., in an unsupervised manner, is a time-consuming process, we applied the proposed
605 criterion to a greedy algorithm, GSSR; to create spectral bands approximating the original
606 reflectance. The final band configurations obtained by GSSR are sets of continuous spectral bands
607 covering the whole spectrum, which preserves the physical meaning of the features like the FS
608 techniques. Compared with the PCFA method providing the best spectral configuration minimizing
609 the representation error, GSSR yields band configurations almost as accurate as PCFA, but in a
610 much faster way. Applying the algorithms to more than 1000 diverse spectra to provide spectral
611 configurations with 5 to 30 bands, GSSR was 50 to 1200 times faster than PCFA, while the mean
612 difference in RMSE was 0.0002 on reflectance scale, i.e. in [0,1]. The difference was even less
613 when both algorithms provided the same spectral configuration for an entire hyperspectral scene,
614 i.e. 0.0001. Using the reduced band configurations in a target detection experiment showed that the
615 bands provided by GSSR are more correctly identified than by PCFA if the spectral configuration

616 has one band more, while the former can be obtained 2000 times faster. The classification overall
 617 accuracy over two hyperspectral datasets using two classifiers revealed that the proposed criterion
 618 provides a band configuration giving a higher classification accuracy than its FS competitors.

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623

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